# caeleste (

A rad-hard, global shutter, true HDR, backside illuminated image sensor

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### Purpose

To realize a family of image sensors particularly suitable for space missions

- 1. BSI (Backside illumination)
- 2. TID (total ionizing dose) radiation-hard design
- 3. "true" or "MAF-" **HDR** (Motion Artifact Free High Dynamic Range)
  - ⇒By reading the same photocharge on two different conversion capacitances
- 4. IWR Global shutter using "**GS**" CMOS technology, which
  - ⇒ Allows low noise readout using CDS (correlated double sampling)
  - ⇒ Enabling Global Shutter without dark current penalty

### In this paper we demonstrate the combined operation of GS and HDR

It is developed under ESA contract 4000116089 "European Low Flux Image Sensor", in collaboration with LFoundry (I) and Airbus (F)

Patent US9819882 granted

## outline

- 1. Purpose
- 2. Principle of operation
- 3. How to create a High DR signal from two "normal" DR signals
- 4. Movie of simultanous GS and HDR operation
- 5. Conclusions

### Key specifications as measured on ELFIS Rev.A

#### Geometry

- ✓ 15µm pixel pitch
- √ 1920x1080 pixels
- ✓ Leftmost 16 columns are test pixels and reference pixels
- √ 16 parallel differential output channels
- ✓ BSI
- ✓ Stitching compatible

### Electrical performance

- √ 40MHz nominal pixel rate per output channel
- √ 500mW at nominal speed

### Electro-optical

- ✓ Two Q<sub>FW</sub> ranges:
  - 1. 10000e (6000e)
  - 2. 160000e in IWR; 320000e in ITR
- ✓ Q<sub>noise</sub> 6.5e<sub>RMS</sub> at nominal speed, excl. DCSN
- ✓ Peak QE > 90%
- ✓ MTF >60% for thin layer BSI
- ✓ PLS >200:1 for thin layer BSI, >500:1 for thick epi\*
- ✓ I<sub>dark</sub> @RT: PPD:60e/s SN:80e/s FD=140e/s

Dynamic range in global shutter, nominal speed

- ✓ ITR, excl DCSN: 160000:6.5 = **87.8dB**
- ✓ IWR, excl DCSN: 320000:6.5 = 93.8dB

<sup>\*</sup> Estimated not measured

2. Principle of operation

### ELFIS HDR concept

The high dynamic range is obtained by converting the photocharge of a given integration time to two signals (voltages).

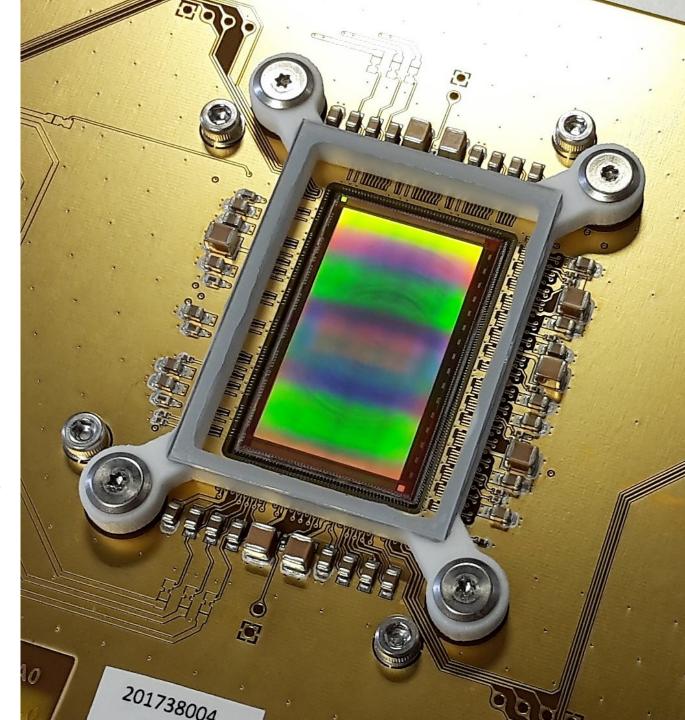
A high and a low conversion capacitance is used, corresponding to two  $Q_{\text{FW}}$ :

- 1. Low Q<sub>FW</sub> (high gain): ~10000e-
- 2. High Q<sub>FW</sub> (low gain): ~320000 e-

The low Q<sub>FW</sub> uses the "classic" **S**torage **N**ode of a GS CMOS technology.

The high  $Q_{FW}$  uses overflow capacitors to store an amount of charge that far exceeds the capacitance of the PPD.

In order to realize IWR (integrate while read) two sets of high  $Q_{FW}$  capacitors are used.



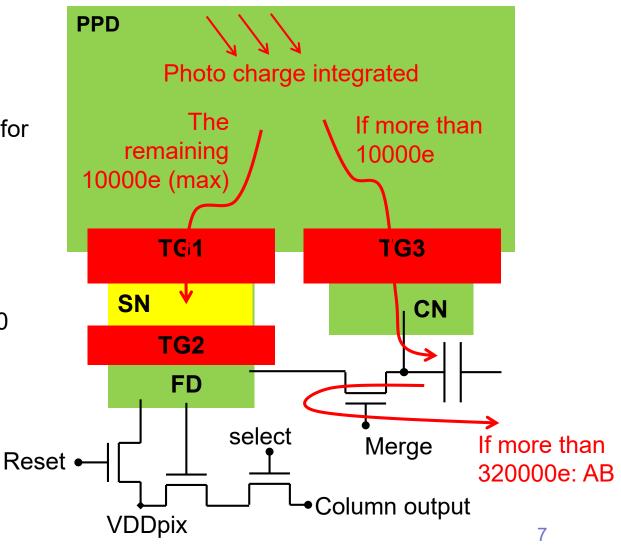
# Operation of the ELFIS pixel: how to read the same photo charge with two conversion gains

During the integration time photo-electrons are accumulated in the pinned photodiode (PPD)

If the amount of electrons in the PPD exceeds 10000, these overflow over TG3 into the capacitor node (CN), for later use. There are two CNs with a total capacity for 320000 electrons

If the total charge exceed 320000 electrons, it goes to the anti-blooming drain.

At the end of the integration time, the (maximally) 10000 electrons still present in the PPD are transferred by transfer gate TG1 to the storage node (SN). The SN cannot contain more.



# Operation of the ELFIS pixel: Caeleste how to read the same photo charge with two conversion gains

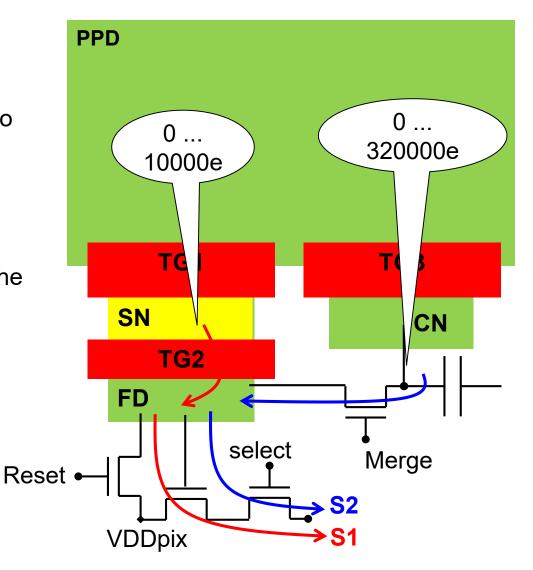
Just before the moment of readout, there are

- Between 0 and 10000 electrons in the SN
- Between 0 and 320000 on the CN(s)
   The total integrated photocharge is the sum of those two

TG2 is toggled and transfers the SN charge to the FD, where is it read out using correlated double sampling (CDS), yielding a signal "S1"

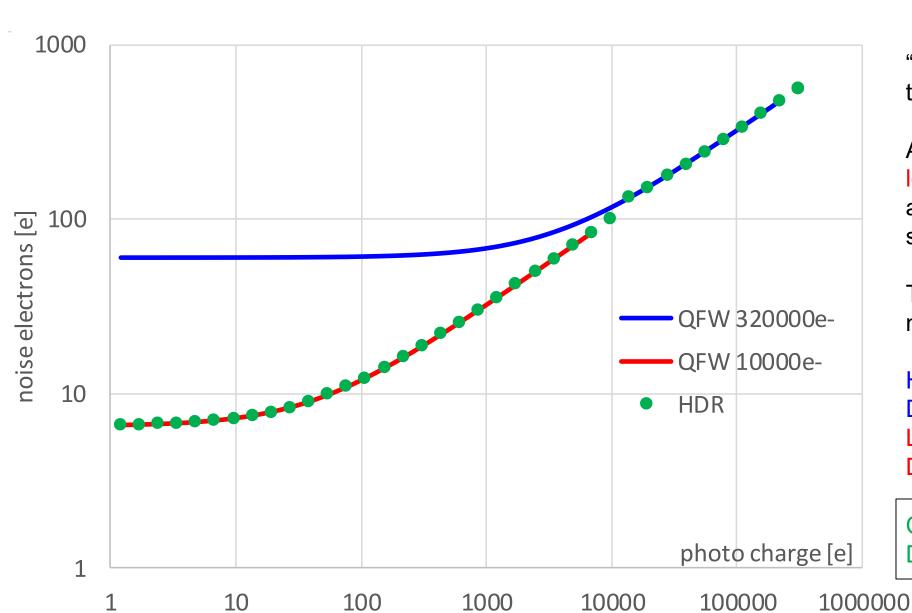
Then "Merge" is closed, shunting the FD and the CN. The sum of both charge packets is on FD, and is read out, yielding a signal "S2"

S1 is the "high gain" signal, with small Q<sub>FW</sub> S2 is the "low gain" signal, with large Q<sub>FW</sub>





### HDR by combining two "normal" DR ranges



"Noise versus Signal" relation of the ELFIS pixel.

At each illumination level both low Q<sub>FW</sub> and high Q<sub>FW</sub> signals are measured. The "HDR" signal is a combination of these

The new dynamic range is the max  $Q_{FW}$  / min  $Q_{noise}$ 

High Q<sub>FW</sub> range:

DR=320000/60 ≈ 5000:1

Low Q<sub>FW</sub> range:

DR= $10000/6,5 \approx 1500:1$ 

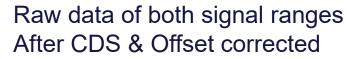
Combination DR=320000/6,5 ≈ **93.8dB** 

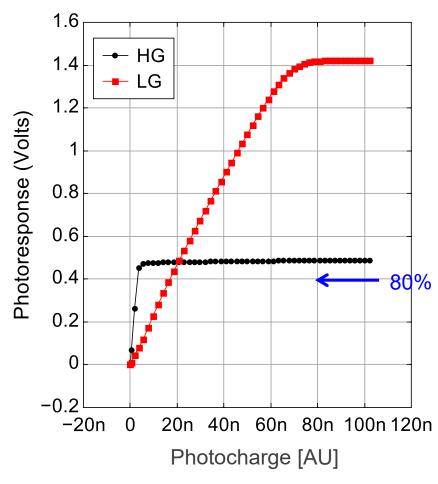
3.

# How to create a High DR signal from two "normal" DR signals

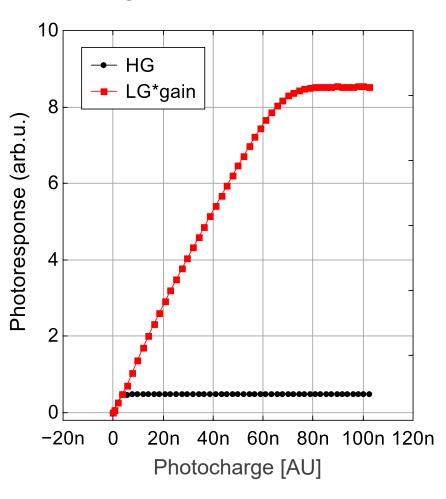
### ELFIS photo response of HG and LG signals







After applying a gain factor on the LG data



ELFIS pixels convert the integrated photocharge almost simultaneously on two conversion capacitances.

These are read out over the same channels time-multiplexed within the row time.

high gain = "low  $Q_{FW}$  range" low gain = "high  $Q_{FW}$  range.

## Avoid discontinuity when merging of HG and LG signals

### smooth interpolation:

When HG signal is above 75% of HG saturation

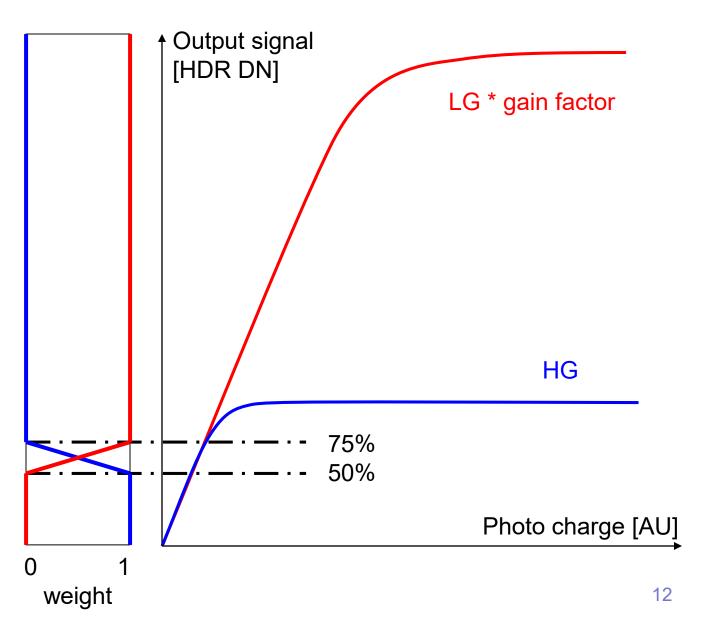
→take the LG signal

When the HG signal is below 50% of HG saturation

 $\rightarrow$  take the HG value.

Between 50% and 75%:

→ apply a weighted average.



**4**.

## **HDR** images and movies

### Setup

#### High dynamic range scene:

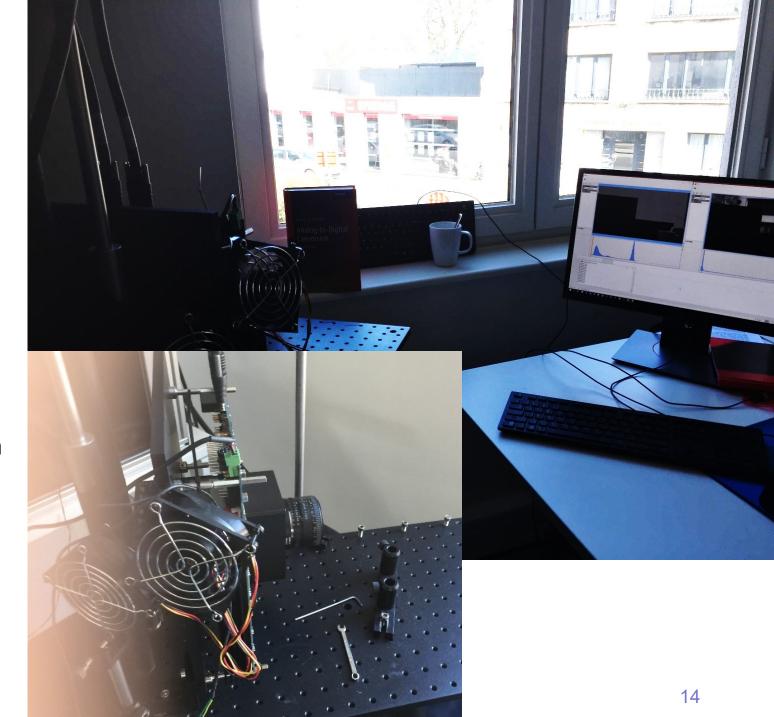
- ✓ Outdoor sunlit buildings
- ✓ Inside lab environment
- ✓ black PC keyboard.

#### The ELFIS breadboard:

- ✓ lens mount with a Nikon 55mm lens
- ✓ Aim for minimal internal reflection in the camera
- ✓ diaphragm set to 22 (! Indeed this 15µm BSI pixel is really light sensitive).

#### Nominal operation conditions

- ✓ IWR (global shutter "integrate while read")
- $\checkmark$  t<sub>frame</sub> = t<sub>int</sub> = 30ms.



## Processing the LG and HG (sub-)frames into a HDR frame Ste





 $(1.977 \times 10^3, 1.599 \times 10^3)$ ; 0.0025

High gain (HG) sub-frame

Here we show the two simultaneous frames of one image, being the HG and LG "subframes" of the same integrated photocharge.

Image processing done here:

In the HG sub-frame on-chip CDS and offchip dark frame subtraction is done. In the LG sub-frame only dark frame subtraction is done.

No PRNU calibration, no linearity correction, no bad pixel or cosmetic corrections are executed.

Low gain (LG) sub-frame

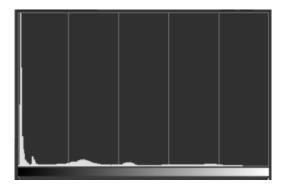
The HDR frame is calculated pixel per pixel as follows:

- → If HG pixel value < 80% of saturation, use that HG value
- → Otherwise use LG \* gain factor

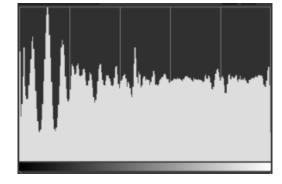
  For demonstration; smoother algorithms exist see further

### Displaying the HDR frame

The resulting HDR image (HDR frame) is stored in floating point format. However, when displaying it by truncation to 8 bit values, the dark parts of the resulting image display as very dark grey or black



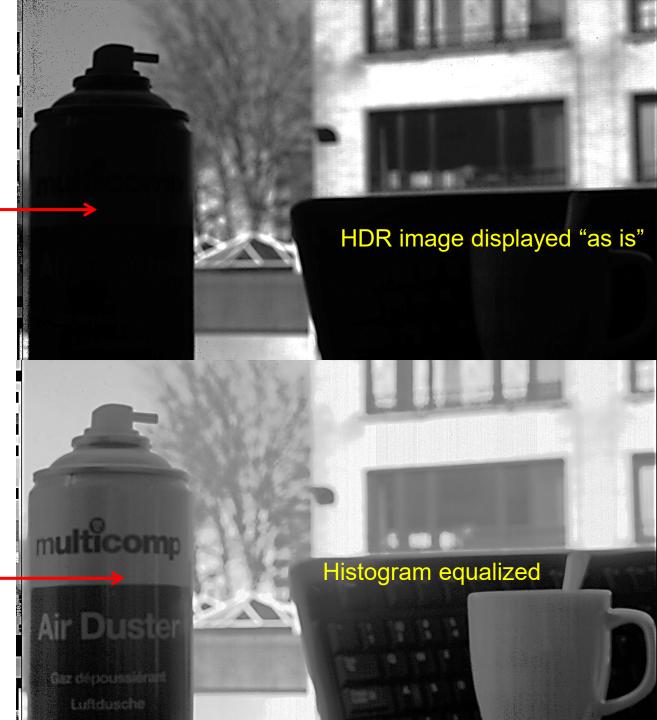
The "raw" histogram



The "equalized" histogram

For display purposes one does "histogram equilization", resulting in a +/- constant distribution of grey values over the image.

Better alternatives exist depending on the purpose and application.



- ✓ IWR  $t_{frame} = t_{int} = 30 \text{ms}$
- ✓ Nikon 28mm lens, diaphragm set to 22 (sic!)
- ✓ CDS & dark frame subtraction
- ✓ No PRNU correction, no linearization.
- ✓ The HDR image created by weighted interpolation between the HG and the LG frames
- ✓ The image is "histogram equalized" for display



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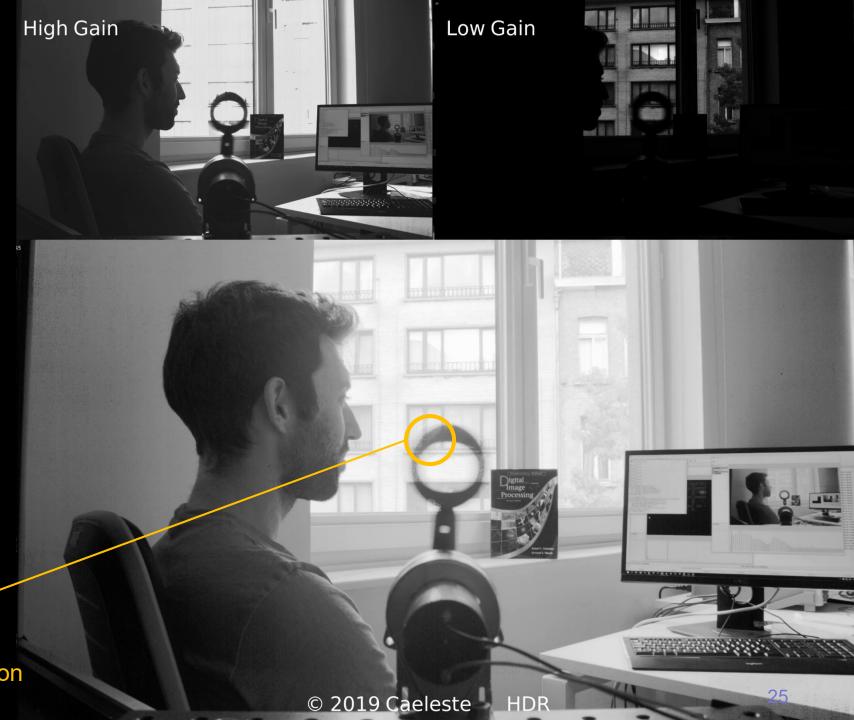


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Artifact is explained by partly illuminated scene non-linearly emphasized by histogram equilization



5.
TAKE HOME MESSAGE

## TAKE HOME MESSAGE

o The combination of true HDR and Global shutter is proven, both as circuit and as technology

- Suitable for imaging high dynamic range scenes, with co-registration of all intensity levels and integration times
  - Star stracker, attitude control, AOCS: global shutter and high dynamic range
  - Earth observations over highly different and unpredicted varying scenes: No gain programming or shutter time adaptation needed by the HDR.
  - Astronomy: high QE by Backside illumination, large range of magnitudes by the HDR concept.